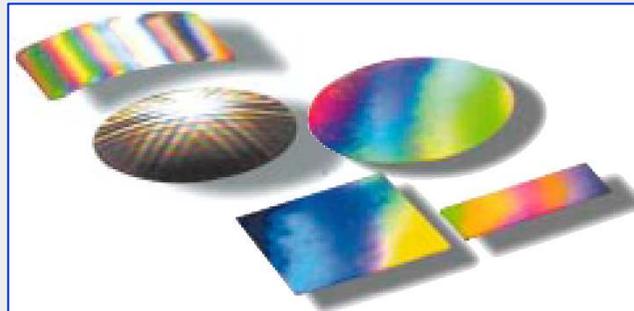




# Square Pore Optics



**PHOTONIS**

**Industrial and Scientific Detectors ISD**



# Square Pore Optics

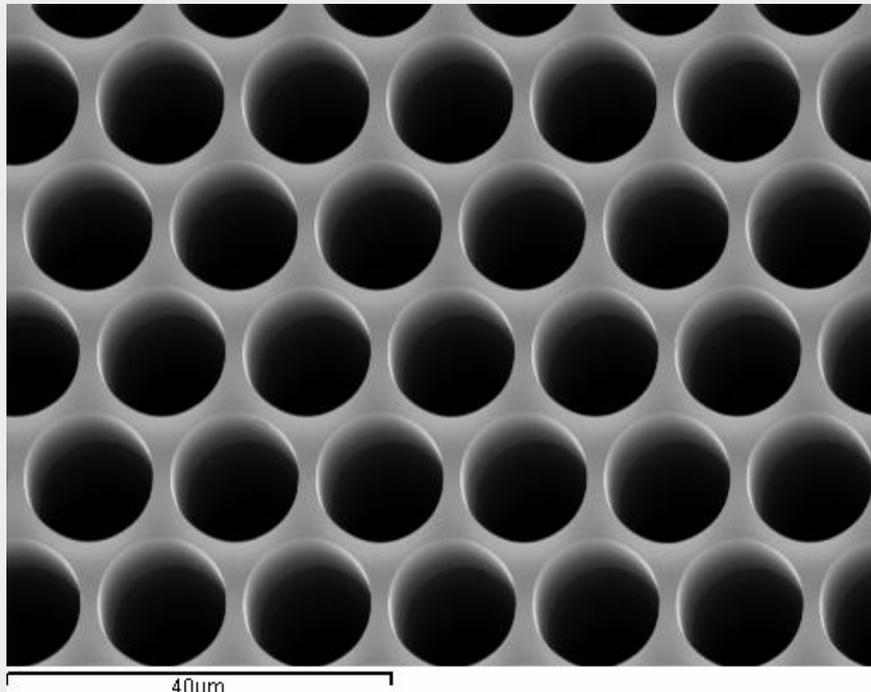
- **Introduction**
- **Production**
- **Applications**
  - **Space: X-ray focussing**
  - **Laboratory**
- **Competitors**
- **Gogogo ...**

# Square Pore Optics



**MPO = Micro Pore Optics (historical name but confusing)**

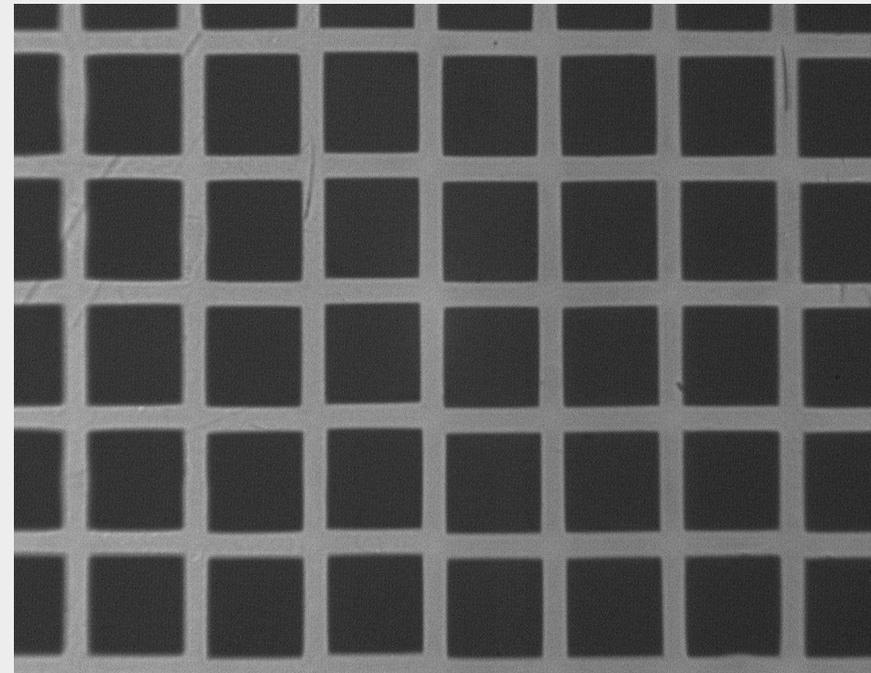
**MCP**



**SEM-picture**

Brive, 12 Dec. 2007

**MPO (20  $\mu\text{m}$  Square Pores)**

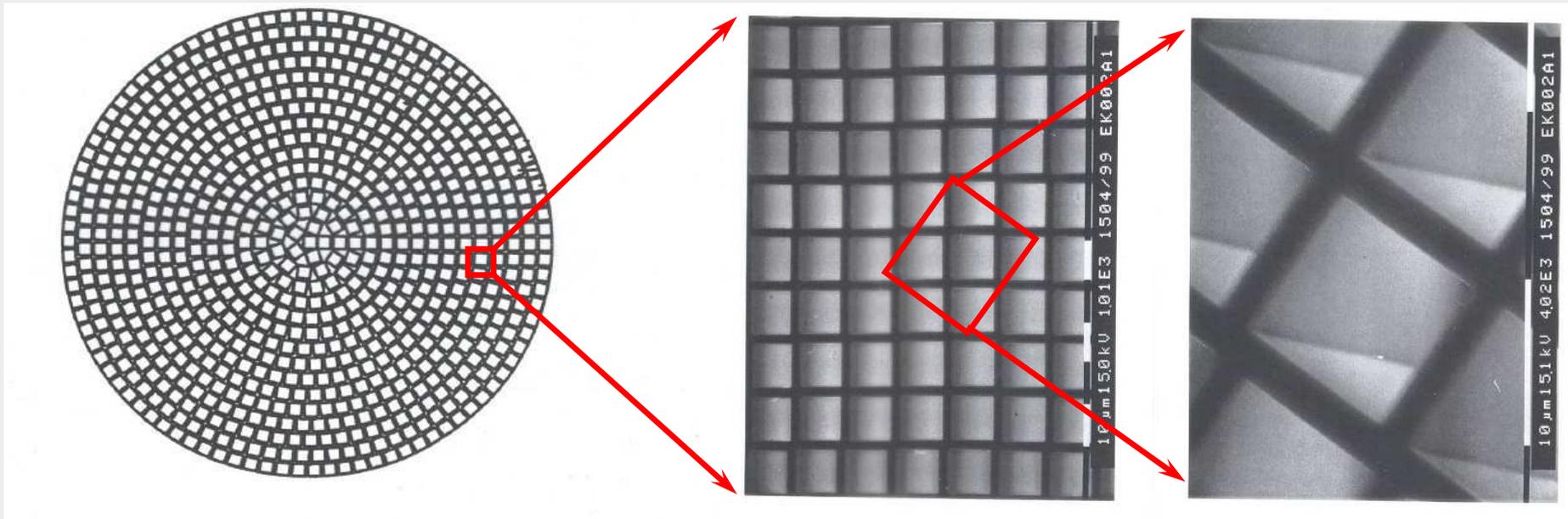


**Optical microscope picture (old)**

E.Schyns, MCP Product Managment



**New technology development by PHOTONIS  
(€ supported by European Space Agency)  
Square Pore Micro Channel Plates**



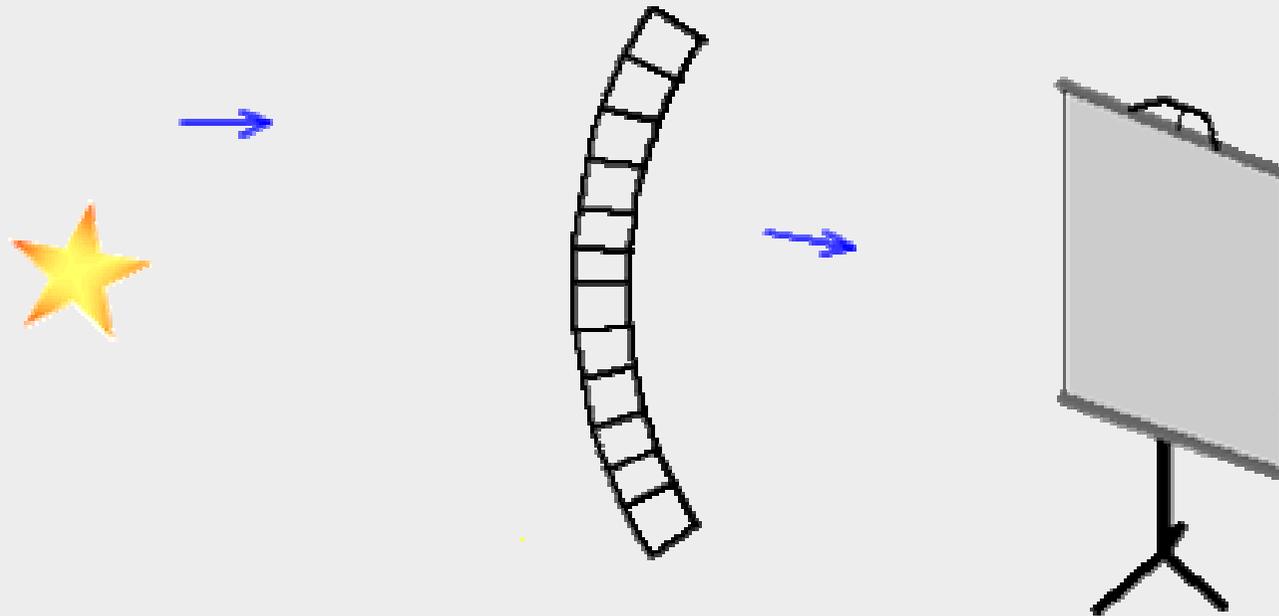
# $\mu$ Pore OptiX : MuPOX



*Distant object*

*Curved ("slumped")  
Microchannel  
plate*

*Screen  
or  
detector*

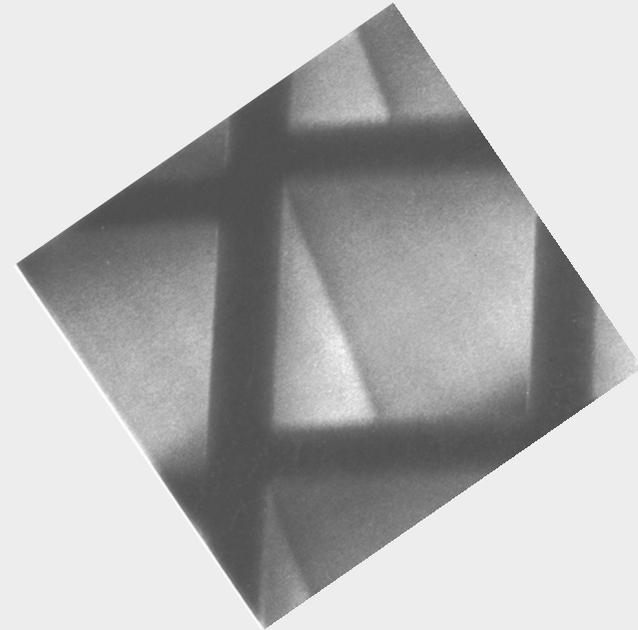
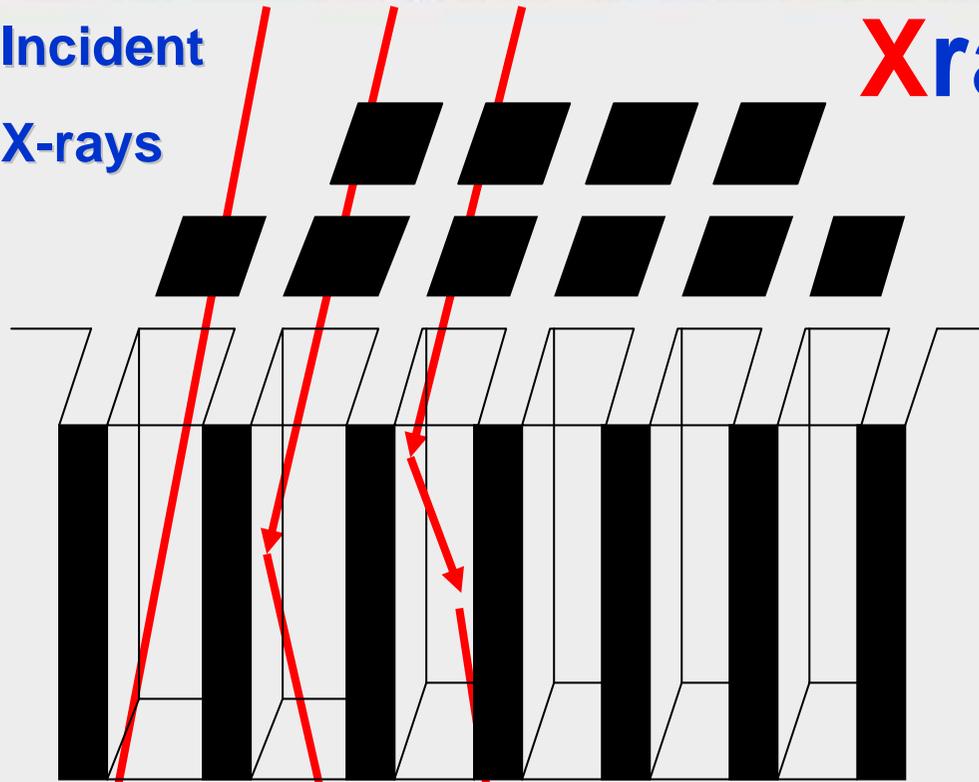


# Operation of an MCP X-ray Lens



Incident  
X-rays

## XrayFLECTOR ! ?



Diffuse  
background

Double reflection (2-axes, true focus)

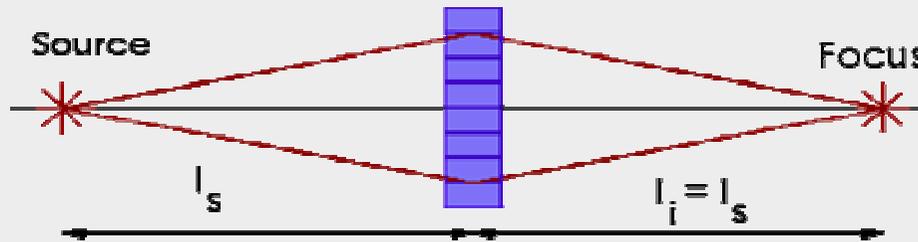
Single reflection (1 axis, line focus)



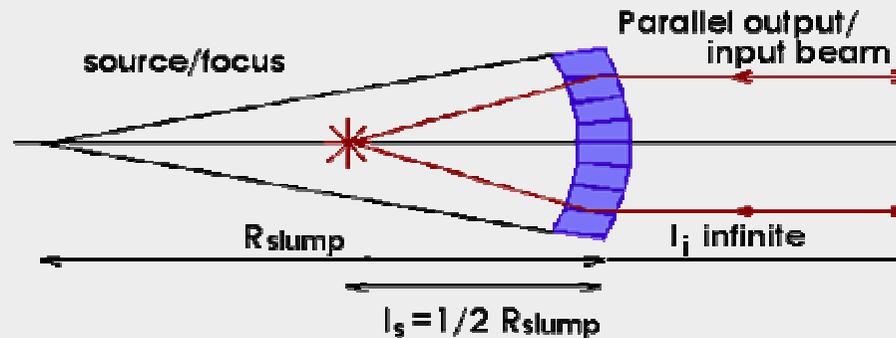
## Focussing / Imaging

**FLAT:**

$$I_{\text{image}} = I_{\text{source}}$$



**SLUMPED:**



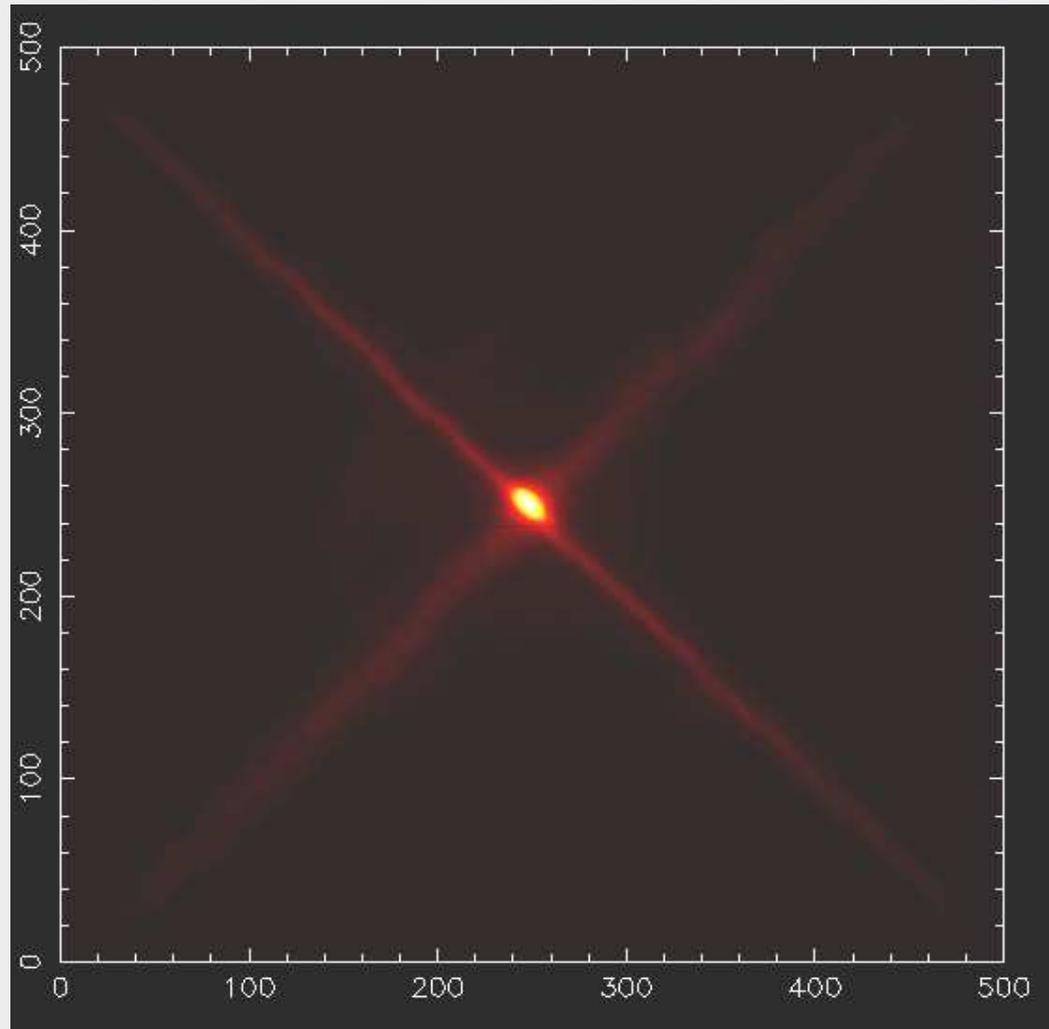
**Make parallel output-beam:**

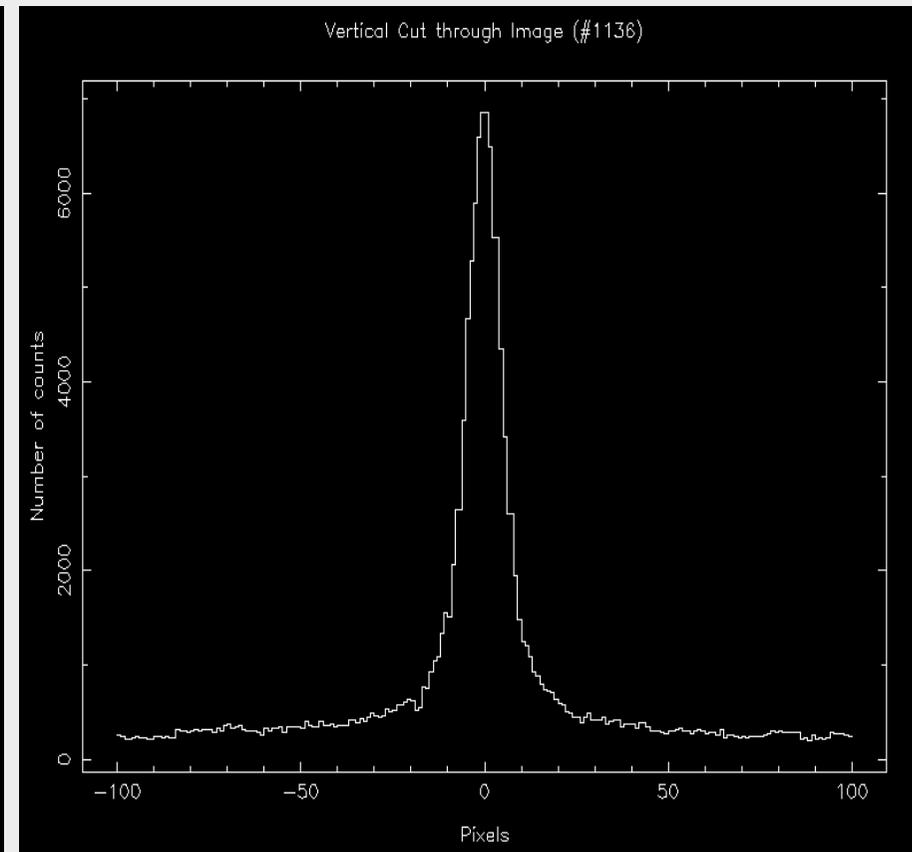
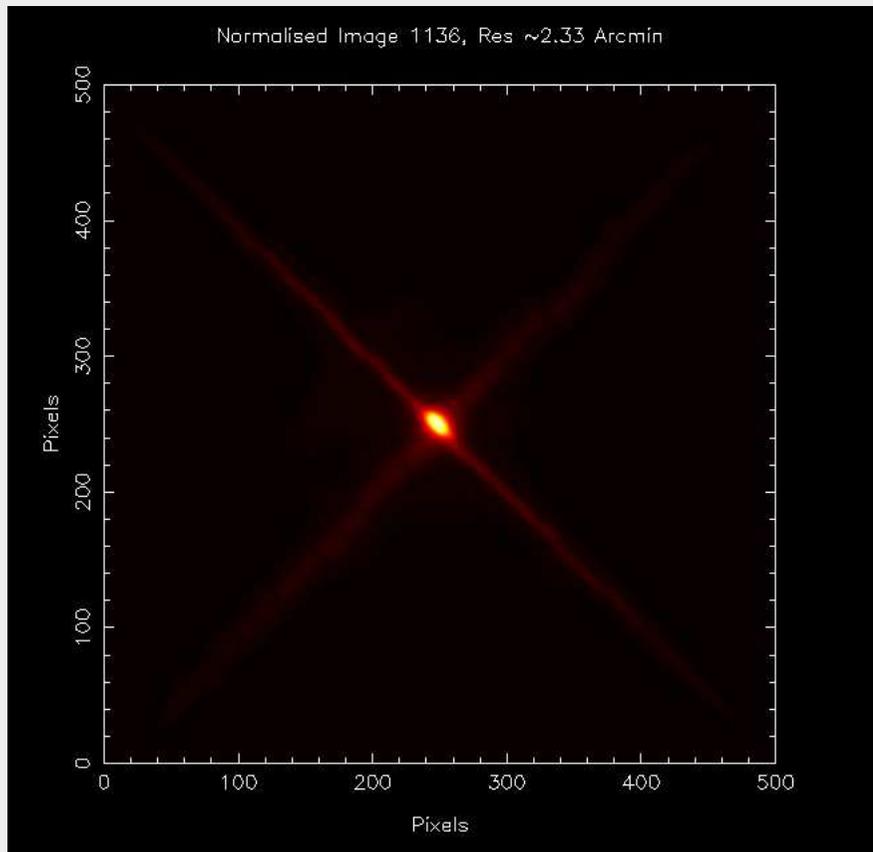
$$I_{\text{source}} = \frac{1}{2} R_{\text{slump}}$$

**Focus (quasi)-parallel beam into a spot:**  $I_{\text{spot}} = \frac{1}{2} R_{\text{slump}}$



## MCP X-ray image quality







**focusing X-rays by reflection off the inside surfaces of square channels**

**Imagine channel plates as collection of millions of little “micro-mirrors” :**

- compact**
- low mass**
- relatively cheap**

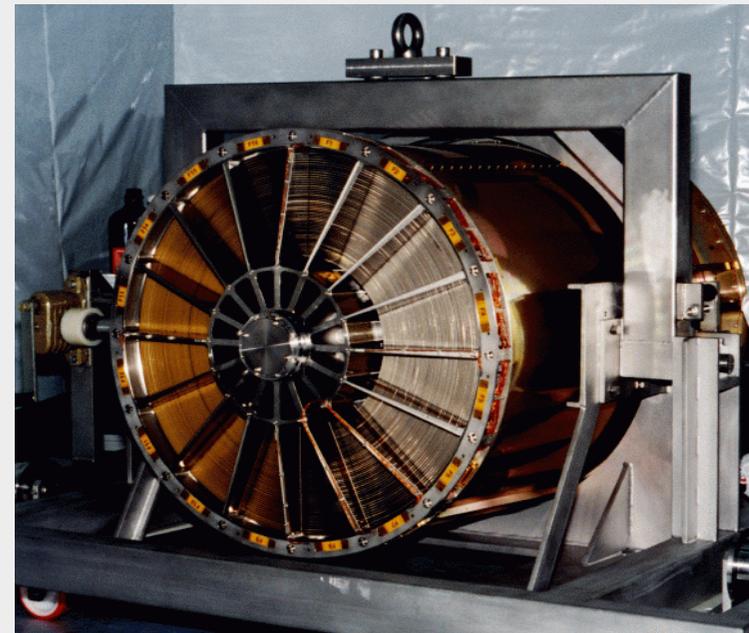
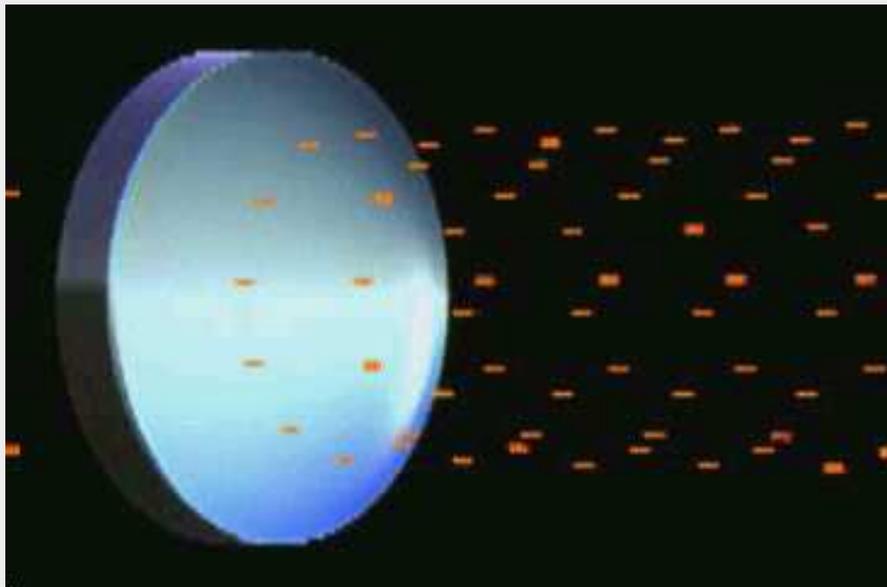
# PHOTONIS Square Pore OptiX: SPOX?



**The old technology:**

**Concentric gold plated cylinders**

**focus / reflect X-rays onto a detector**



**Mirrors for XMM Space project**



## NEW vs OLD TECHNOLOGY

	<b>XMM</b>	<b>SPO</b>
<b>WEIGHT</b>	<b>350 kg</b>	<b>29 g</b>
<b>DIAMETER</b>	<b>700 mm</b>	<b>60 mm</b>
<b>Platform</b>	<b>910 Kg/m<sup>2</sup></b>	<b>10 Kg/m<sup>2</sup></b>

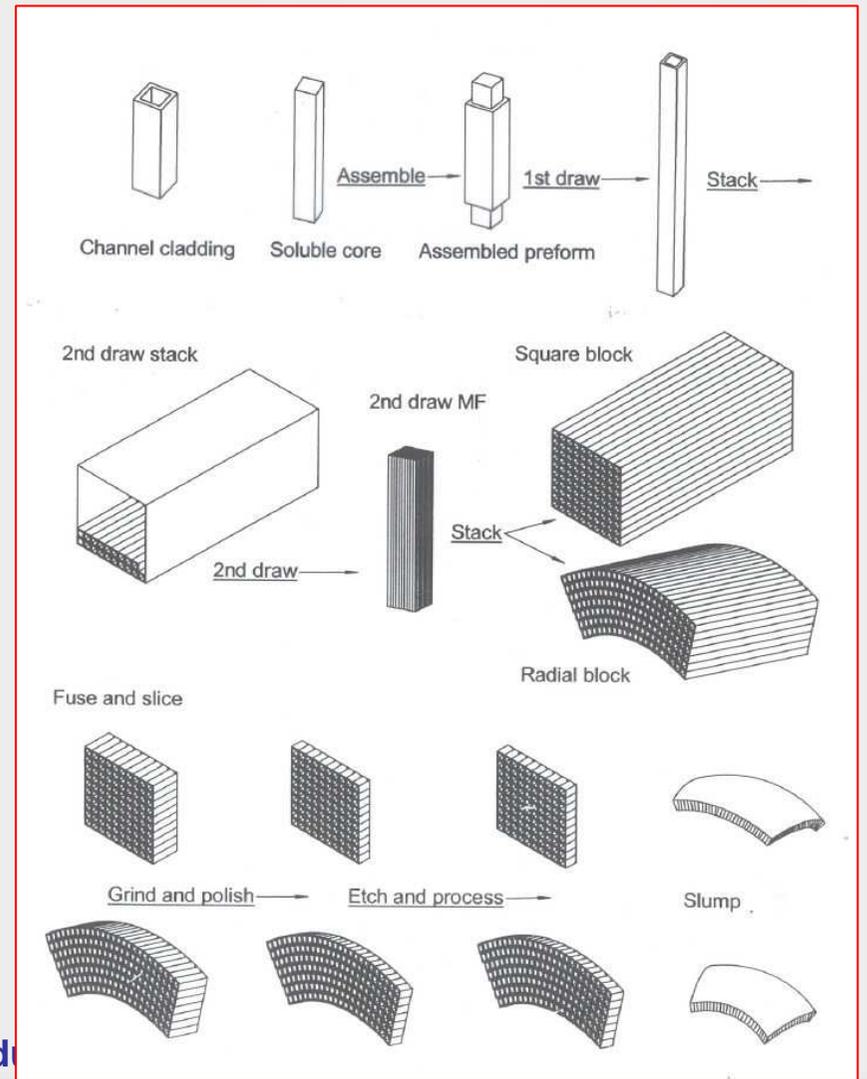


## Square-Pore-Square-Pack (Square MF stacking)

Or

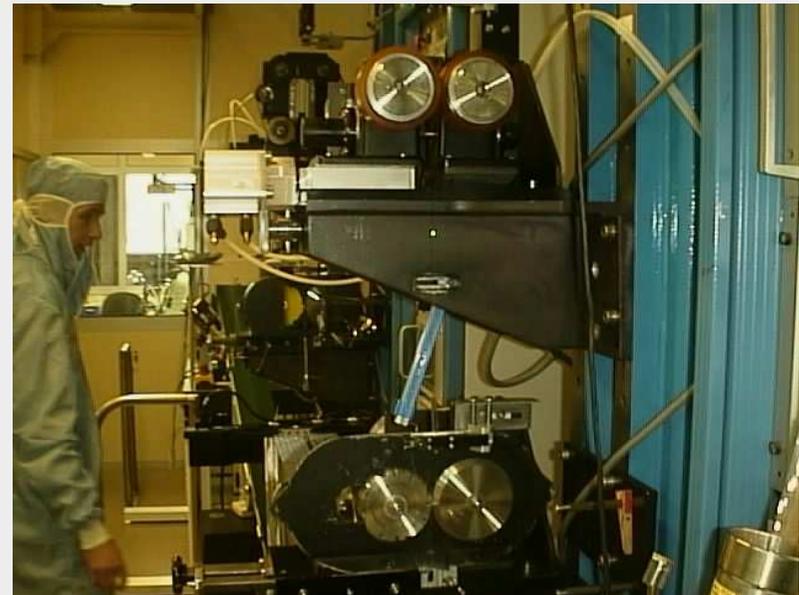
## Square-Pore-Radial-Pack (Radial MF stacking)

More difficult to make (>€) but  
better optical (focal) properties





**drawing tower  
automatic fibre  
or  
multifibre “kickers”**



**Automatic fibre and  
multi-fibre monitor**



**Square stacking of square multifibres (square-pore square-pack)**



**Stacking with video and software for quality control**



At high photon energies i.e. very short  $\lambda$ , scattering from surface roughness is an important source of loss

**Currently we make channel roughness ~5 nm rms**

## ETCHING CORE TECHNOLOGY

- need small angles of incidence ( $< 2^\circ$ )
- allow the rays to strike (graze) the surface
- the channels have to be long

NEED HIGH L/D RATIO i.e. the larger L/D the better

(also allowing X-ray FLECTION with higher energies ...)



**The resolution will be mainly due to:**

**The geometrical quality of the OptiXs**

**The roughness of the channel's inner surface**

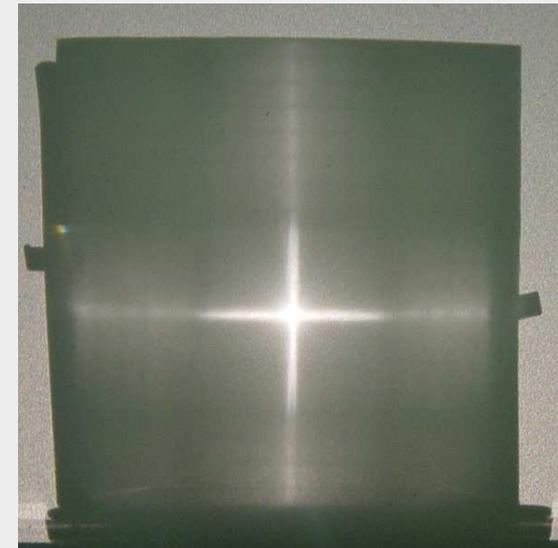
## **Main parameters**

- **Size of the holes (pore size, length, L/D)**
- **Geometry of the channels (fibers, multi-fibers)**
- **Stacking, channel's alignment**
- **Material reflectivity (special glass)**
- **Inner surface roughness**



## Applications

- Scientific / Space applications
- Laboratory applications
- X-ray lithography



(image from visible light)

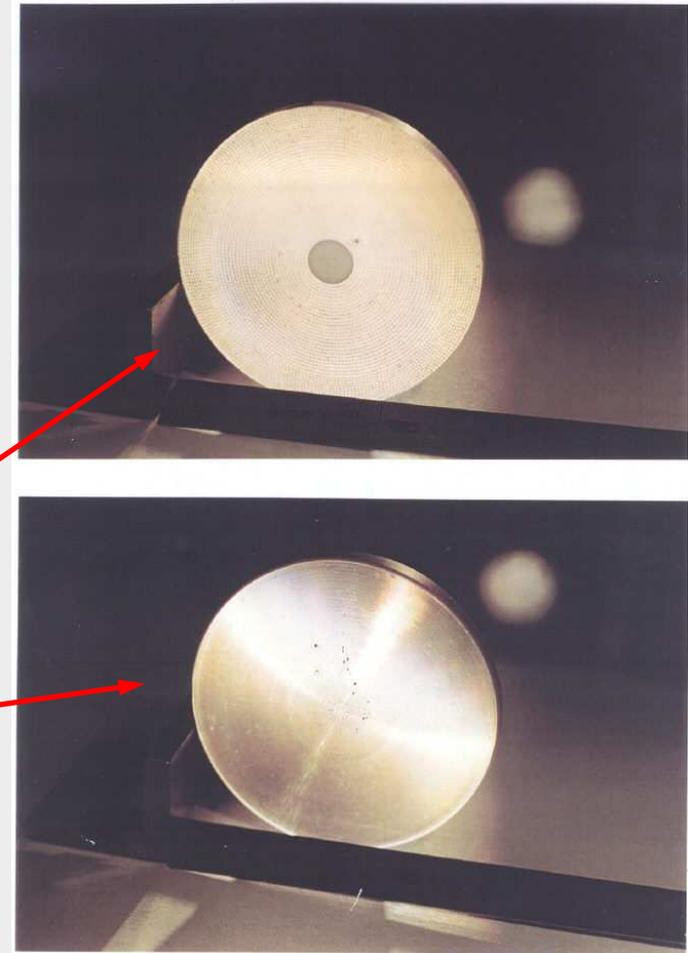


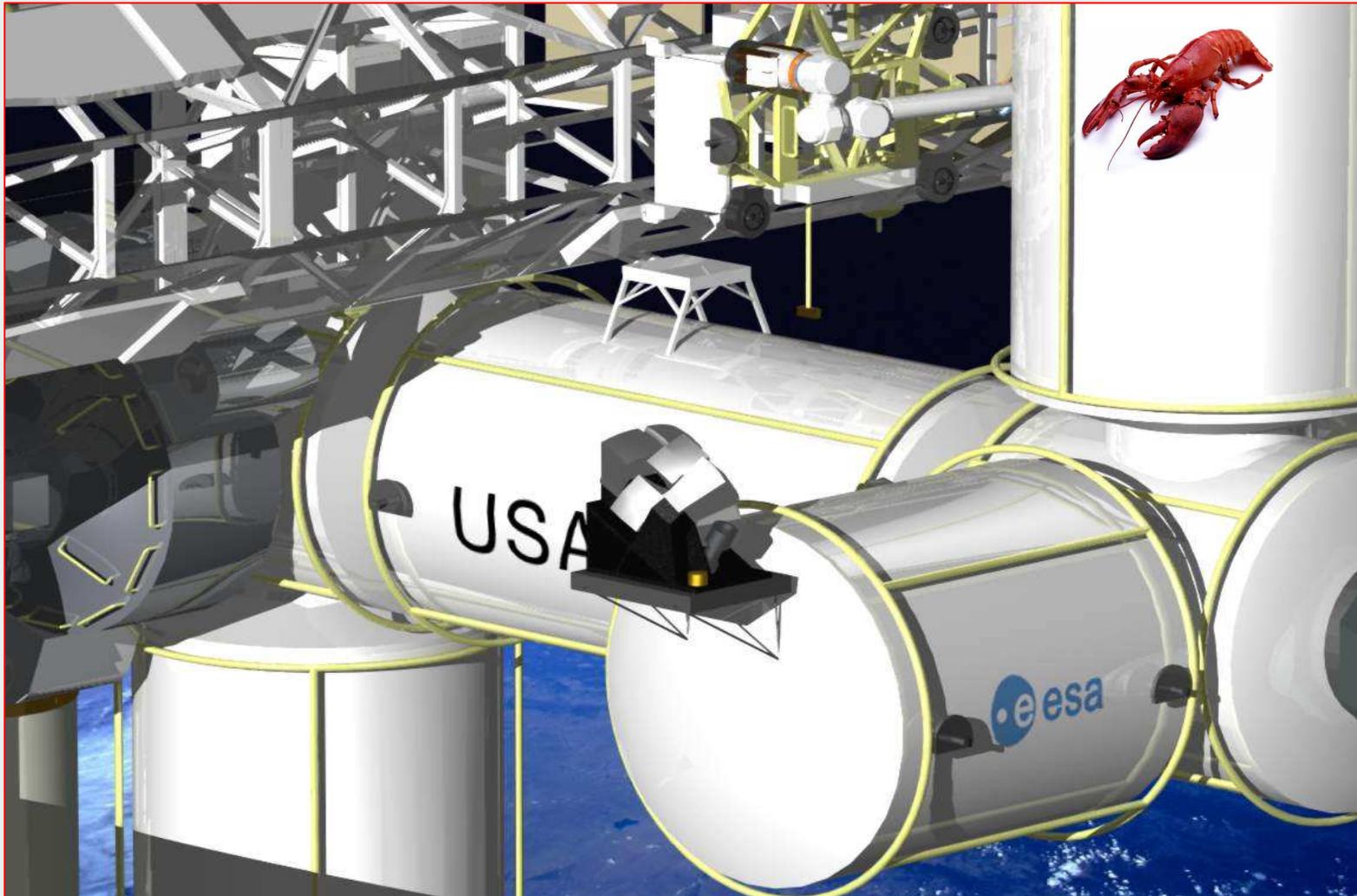
## SCIENTIFIC APPLICATIONS:

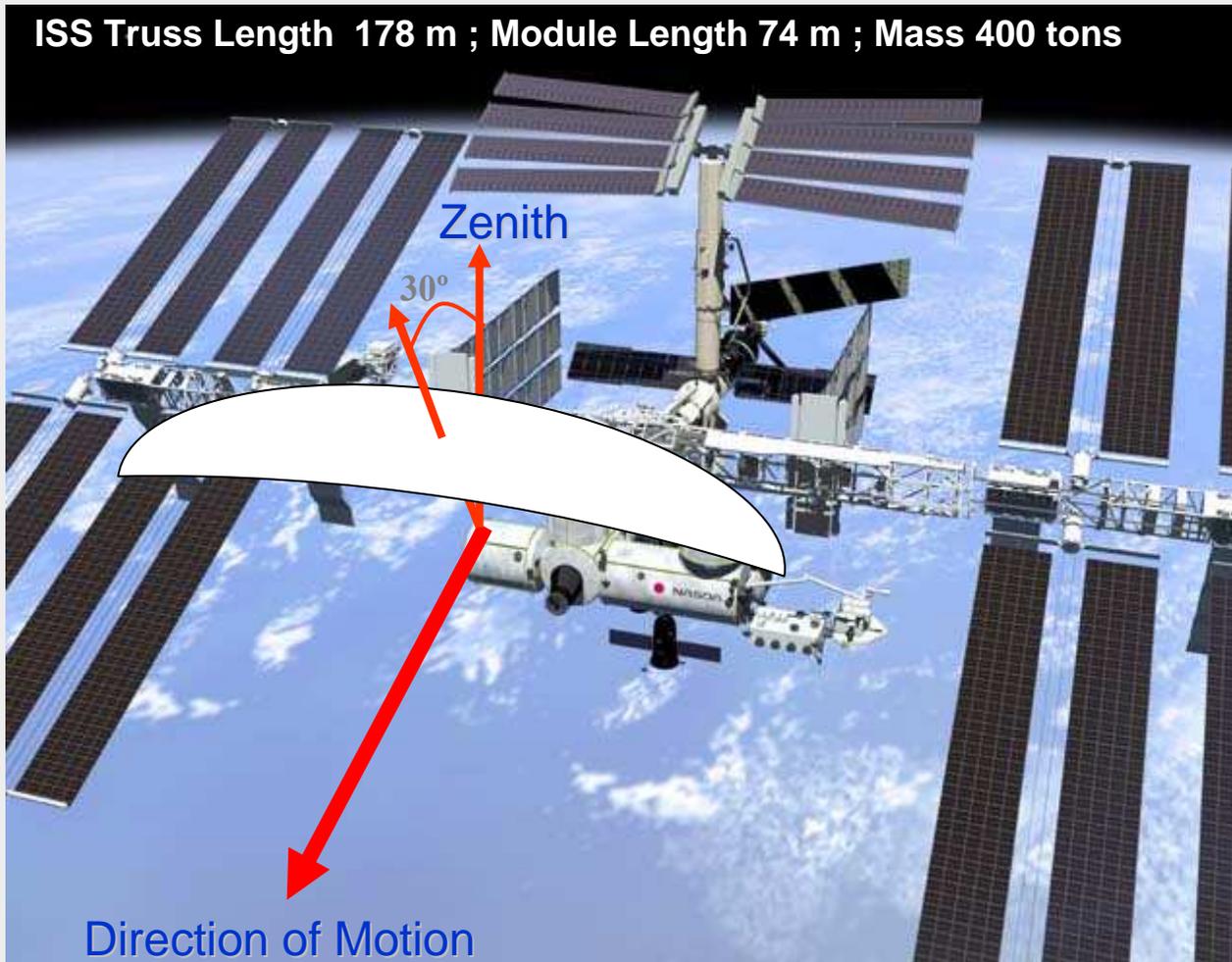
### SPACE

- X-ray domain
- Deep UV (XUV / EUV)

**OptiX made by PHOTONIS**  
**(ESA / ESTEC contract)**







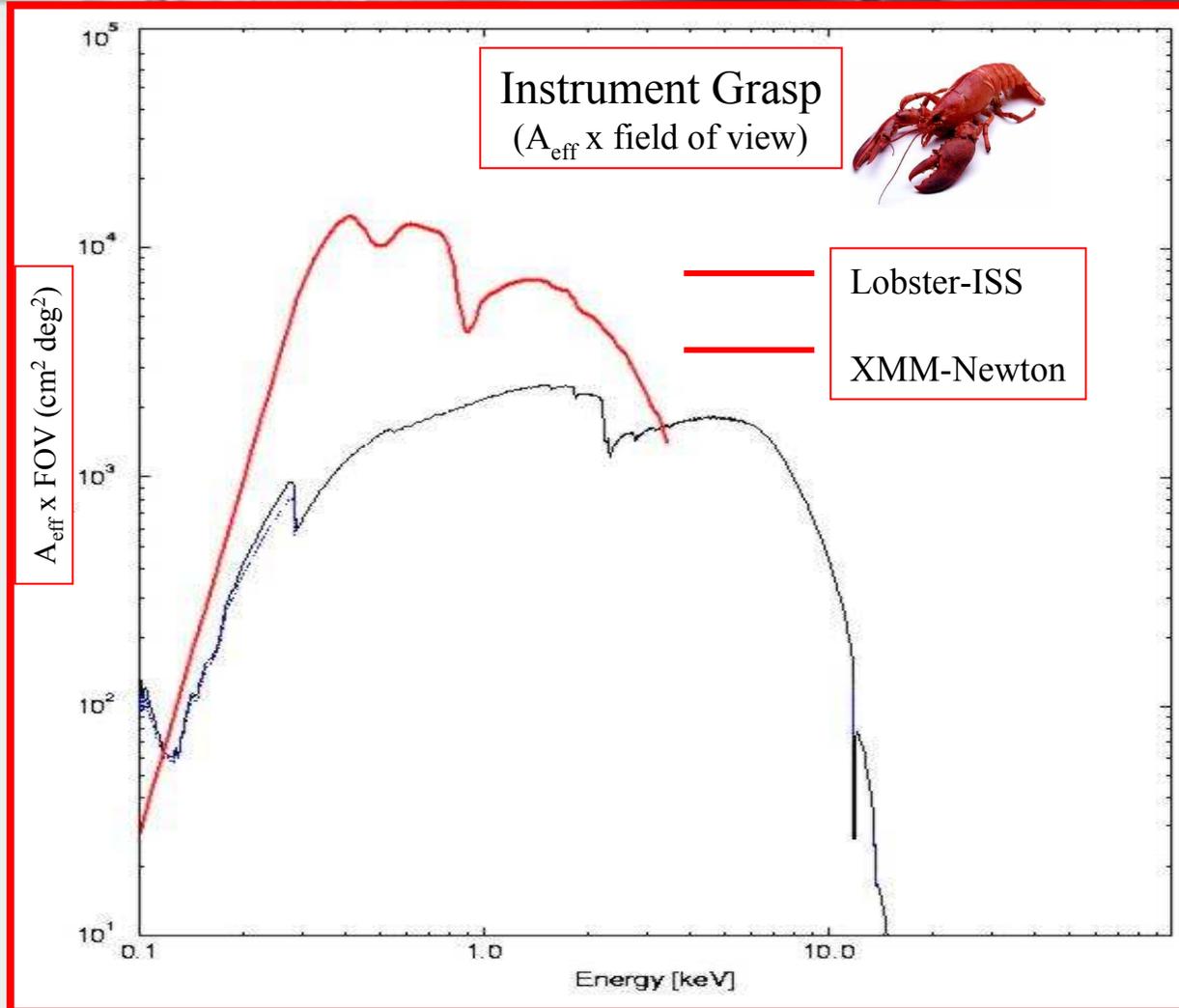
ISS Truss Length 178 m ; Module Length 74 m ; Mass 400 tons

Tilt: 30° forward of zenith  
Total field: 162° x 22.5°

Field sweeps sky every  
90 minute orbit.

Rotation Axis

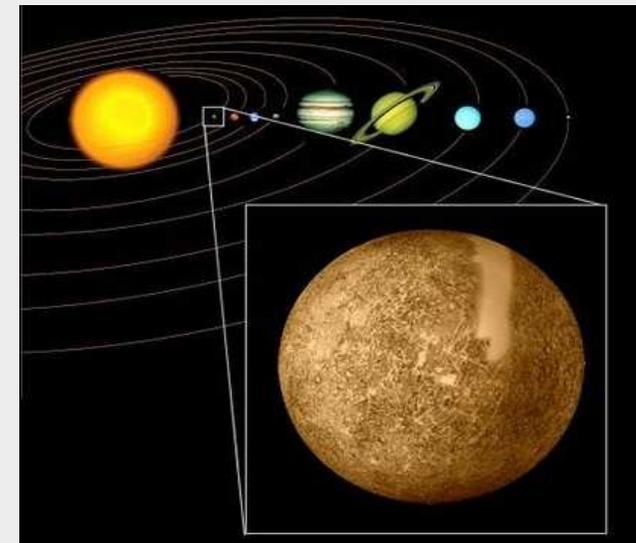
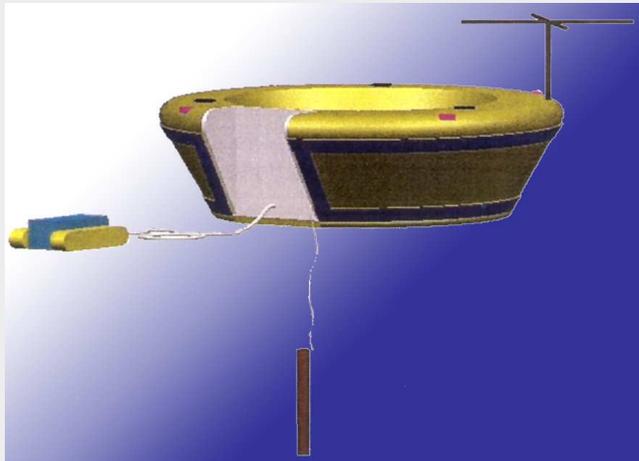
# PHOTONS and XMM-Newton : The impact of novel, low-mass technology





### TARGET: Analyses of Mercury surface (X-ray spectroscopy)

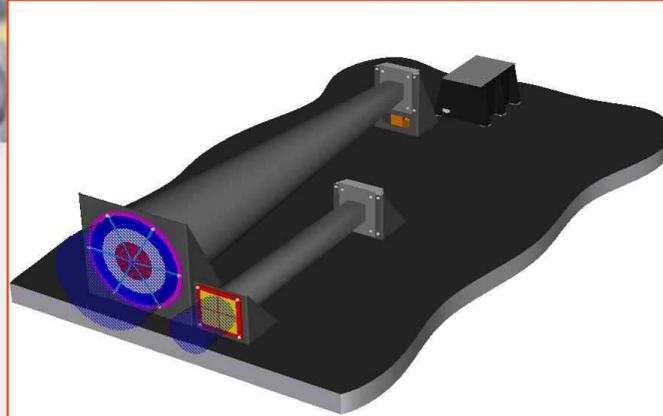
- \* Why close to the sun ?
- \* Why core is big (80% of the mass) ?
- \* Why Mercury has a magnetic field ?
- \* Ice on the pole of Mercury ?



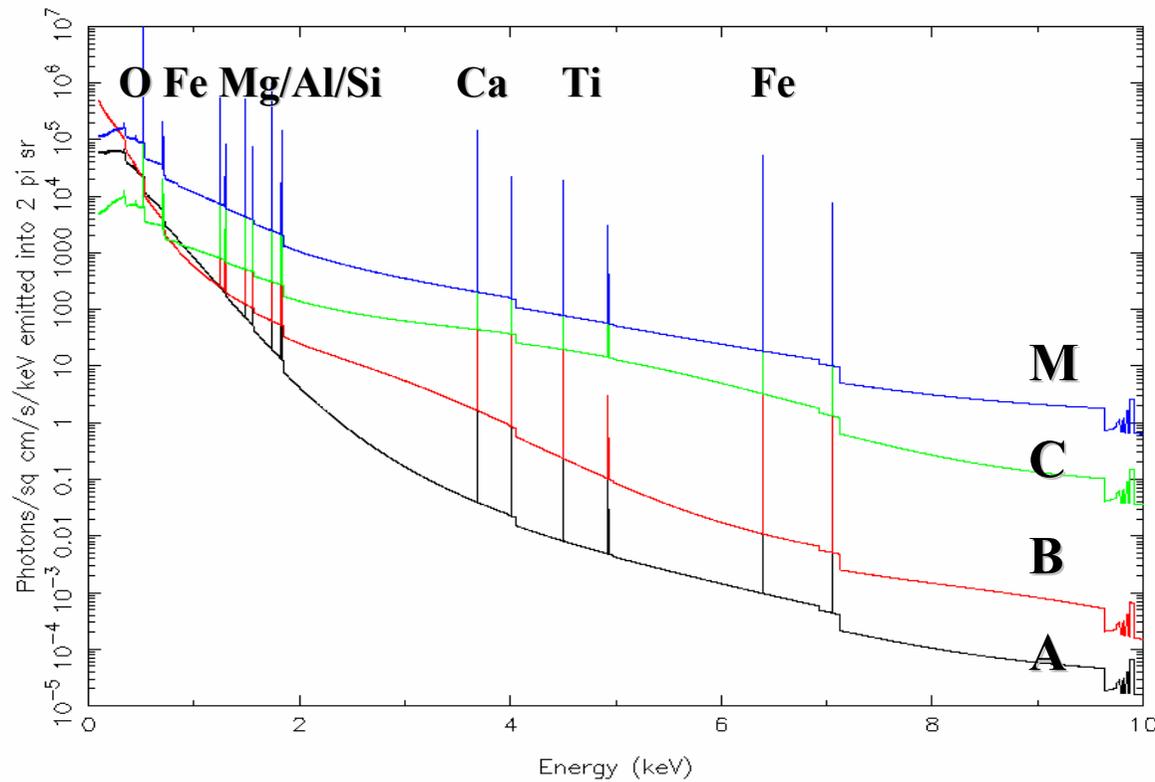
Mercury

On it's way to Mercury, passes by the Moon & Jupiter: switch on?

# Square Pore $\mu$ -OptiX



2 instruments  
- MIXS-T  
- MIXS-C

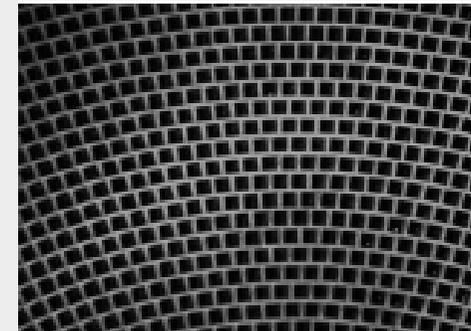
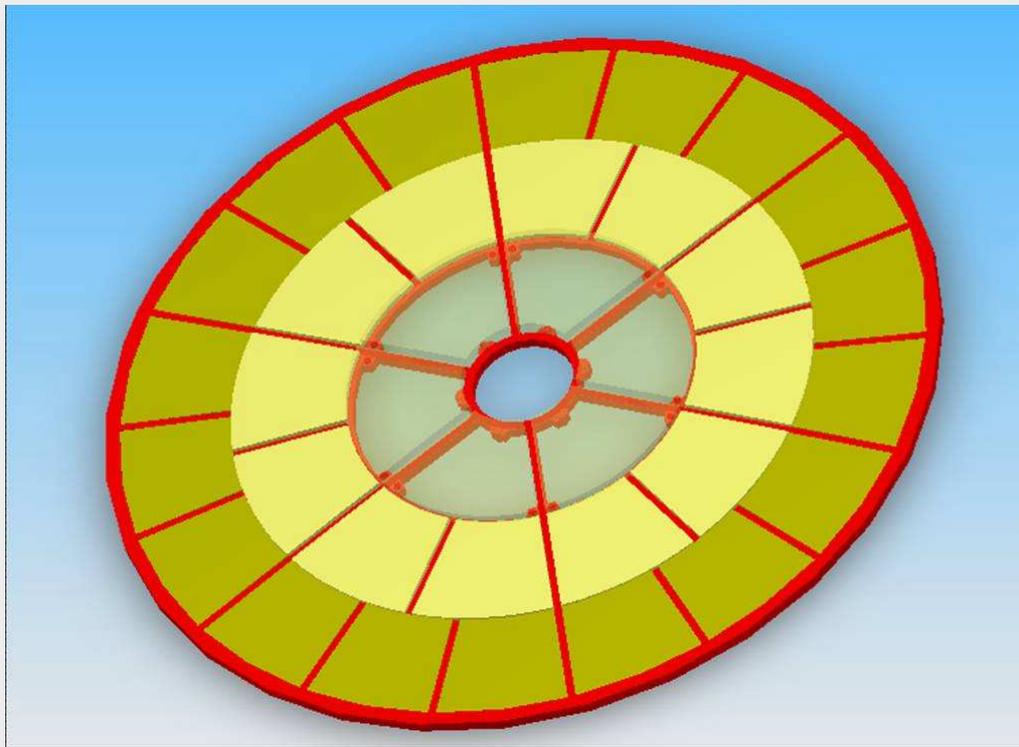




## Detailed MIXS-T optic design

Input to Breadboard development project

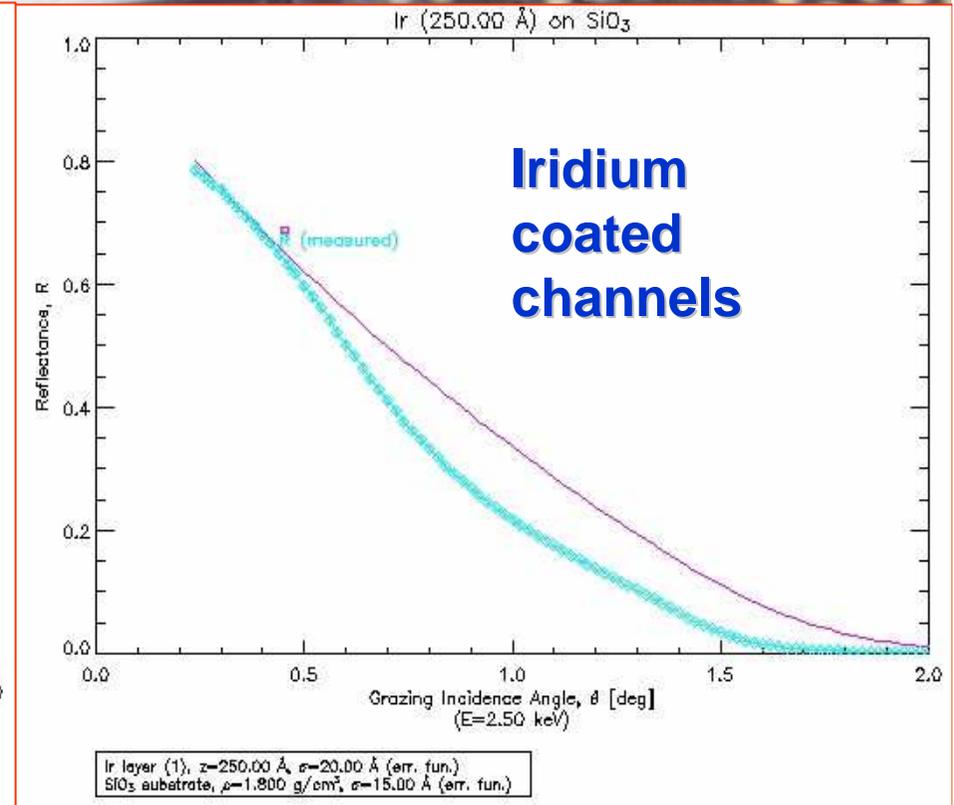
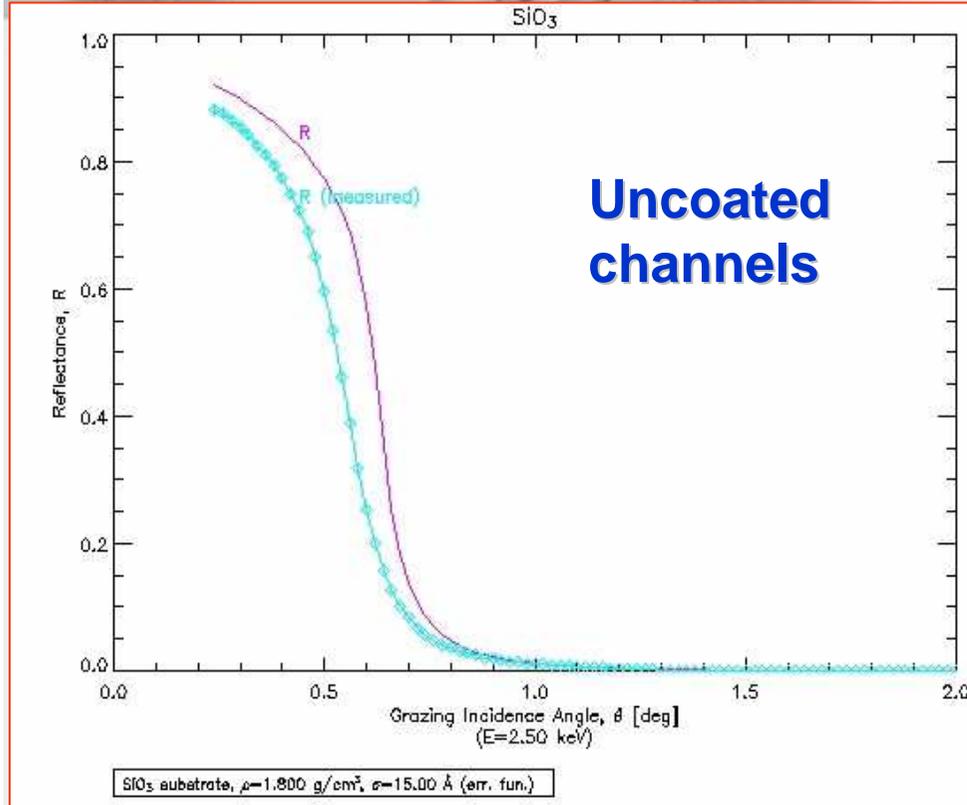
ESA / Photonis / cosine-Research collaboration



6 sectors of  
1-2-3 doublets

$$R_{\text{slump } 1} = 4 \text{ m}$$

$$R_{\text{slump } 2} = 1.3 \text{ m}$$



**Demonstration of enhanced X-ray reflectivity at 2.5 keV for Iridium-coated MCP**

**Other coatings possible: Au, Ni**

**(difficult to get metal all the way in the channels)**

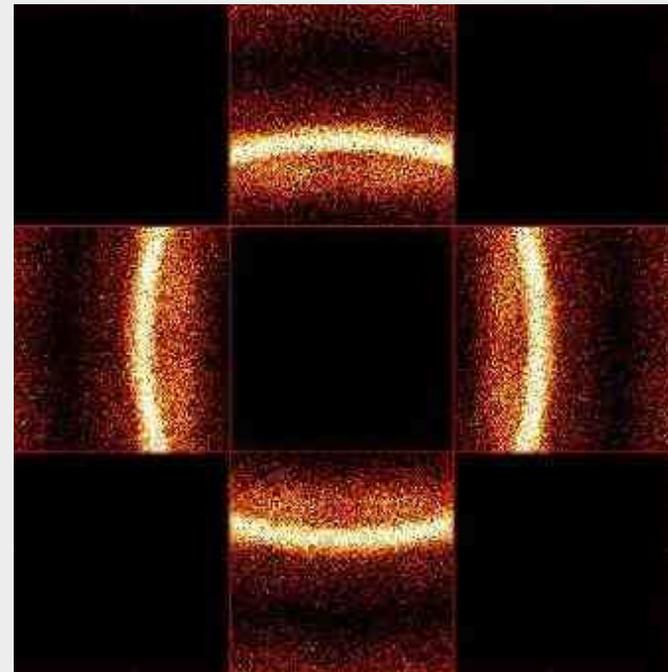
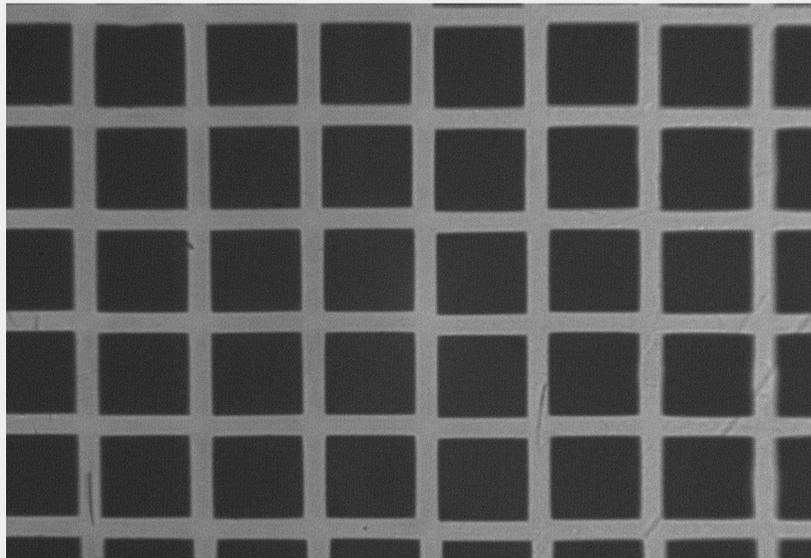


## Mission Auroral Imager (ESA)

Idea from University of Leicester to measure the interactions between particle flux and earth magnetic field at the poles.

Same optics as Lobster programme.

TELESCOP X with 5 modules





## Applications In the laboratory

- Material test and quality control
- X-ray lithography
- XUV lithography
- Medical imaging (decreased dose) !





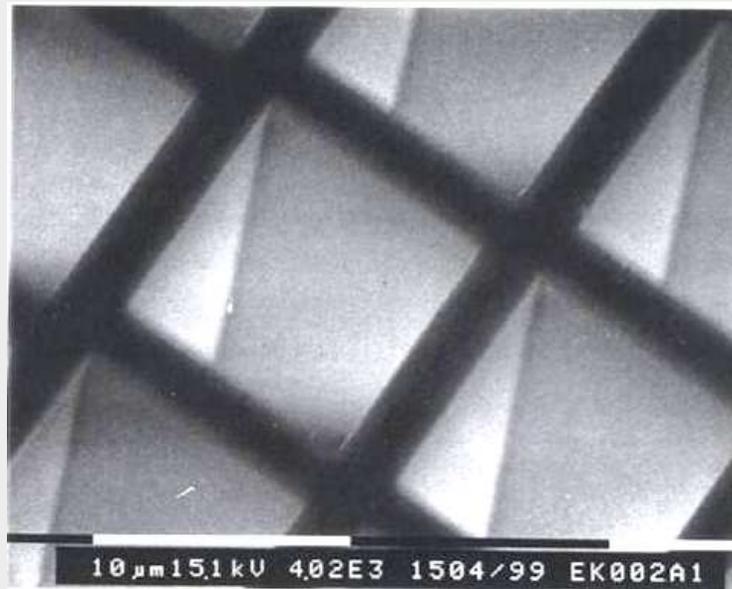
## Applications e.g. X-ray lithography

- The resolution of features on a chip is limited by diffraction from the mask “blurring” the image.
- Diffraction is inversely proportional to  $\lambda$ ,
- Soft X-rays  $\lambda \sim 1 \text{ nm}$  : no problem
- OptiX turns a diverging X-ray beam from a point-like source into a parallel beam and focuses down to a tiny point

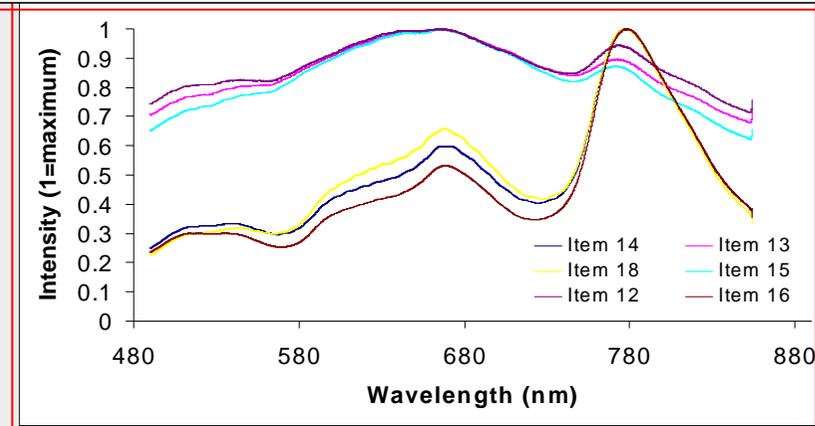
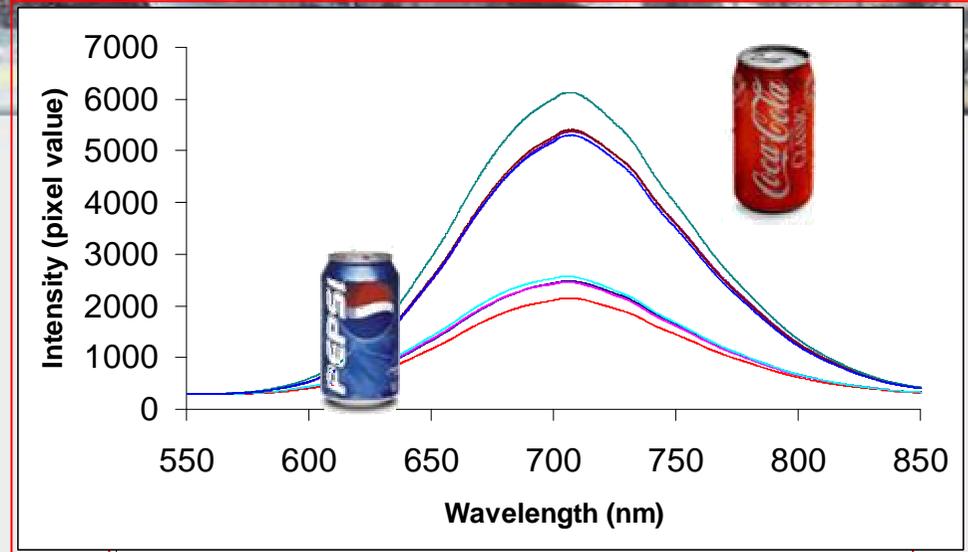
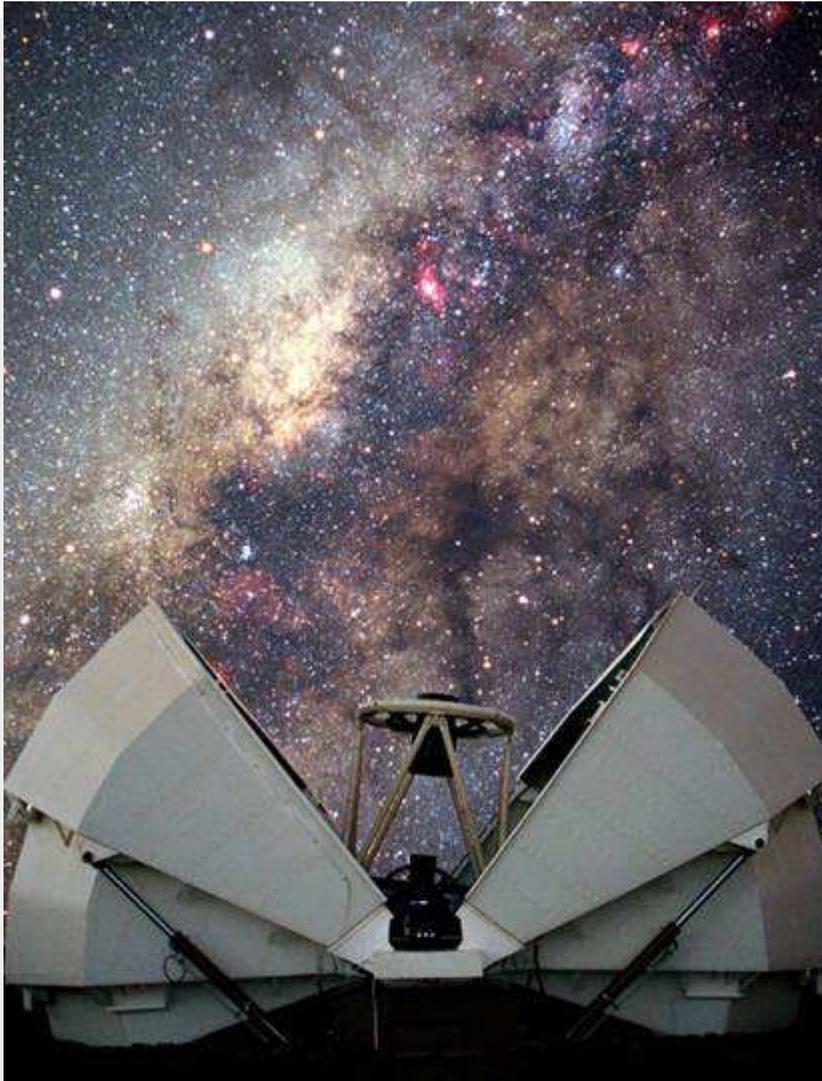


**PHOTONIS is the sole provider, i.e.**

- can produce large quantities**
- special glass composition**



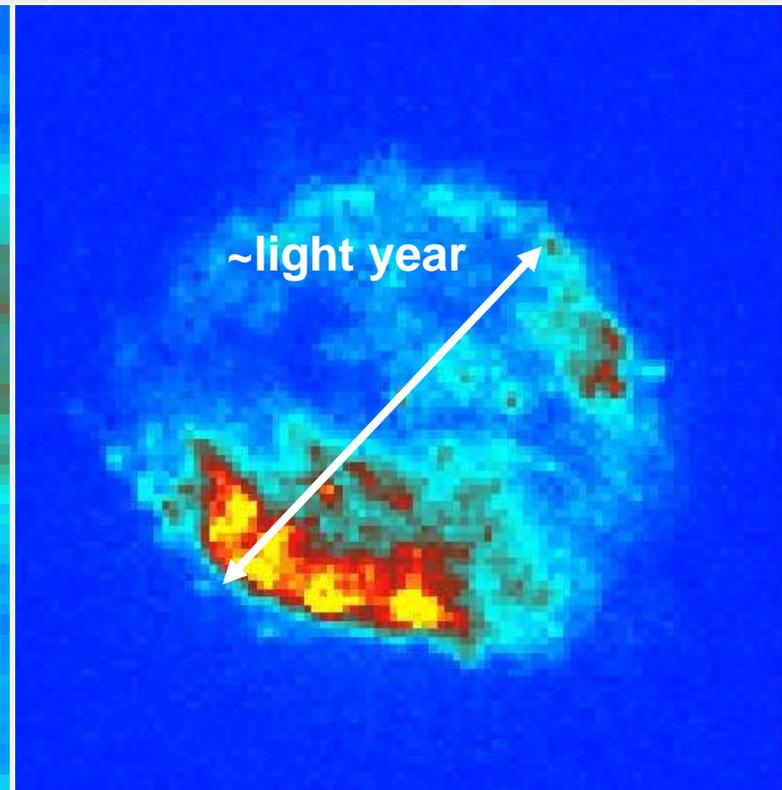
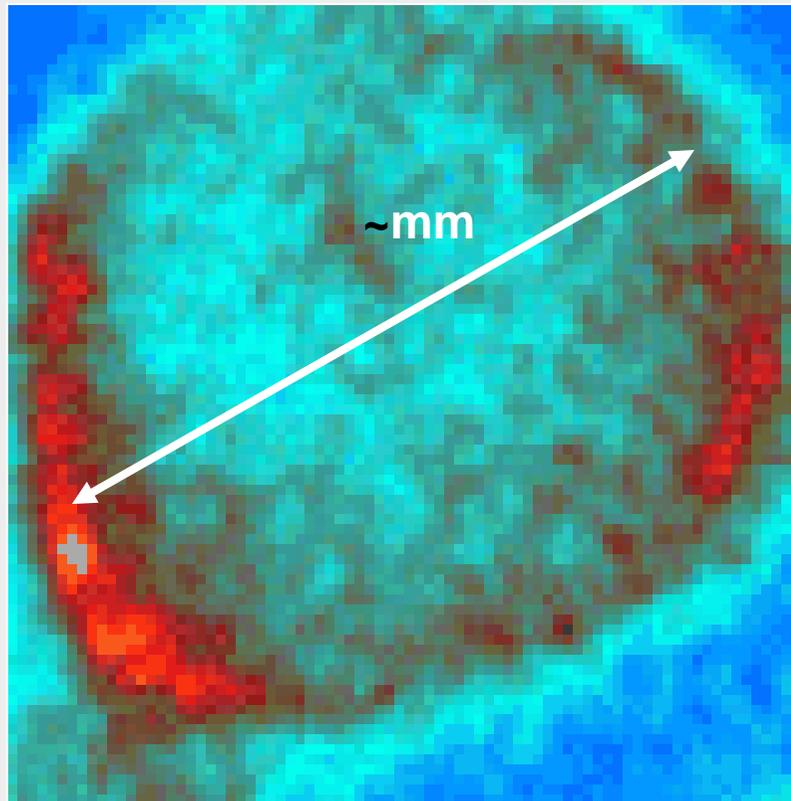
Other examples of applications coming from (outer) space:



**Faulkes Spectrograph –**  
measures the precise colours of objects  
(originally stars) to a high level of precision



**Goals - large area, high resolution, high efficiency, low background**



**BIOIMAGING UNIT**